



PHYS 111 5(4-2) PHYSICS I (MECHANICS)

Chapter 6 Gravitation

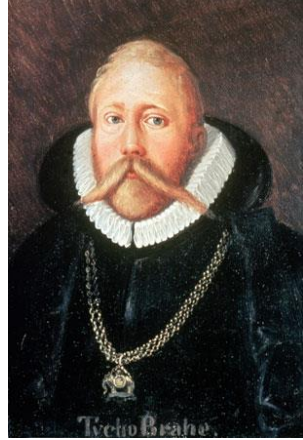
Ahmet Bingöl
METU, Nov 2024



Content

1. Brahe & Kepler
2. Kepler's Laws of Planetary Motion
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Tycho Brahe (1546 – 1601)



- Danish Astronomer who made observations on Planets and Stars.
- He studied law for a while until 1565.
- In 1571 he discovered a bright star in the constellation Cassiopeia and made precise measurements of its position. This star gradually faded away after weeks and gradually disappeared from the sky (is this a supernova?)
- The Danish king gave him a fully equipped observatory to continue his work. Location: Ven island in the Baltic sea.
- On the island, using a Sextant (protractor), he measured the positions of 777 stars and planets and prepared a catalogue. There is no error of more than 1-2 minutes in its measurements.
- Johannes Kepler served as Brahe's assistant for a long time before his death. When Brahe died, he left his entire inheritance to him.



TYCHONIS BRAHE DANI
SCRIPTA ASTRONOMICA

EDIDIT

I. L. E. DREYER

AUXILIO IOANNIS RÆDER

SUMPTUS FECIT G. A. HAGEMANN

TOMUS II



HAUNIÆ MCMXV
IN LIBRARIA GYLDENDALIANA
TYPIS NIELSEN & LYDICHE (AXEL SIMMELKJÆR)

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ÆQUALIS MOTVS LONGITVDINIS SOLIS																	
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	S.	G.	/														
Ianuarius	1	0	33	18	2	1	58	8	1	2	28	0	0	0	30	1	14
Februarius	1	28	9	11	3	2	57	25	3	7	24	2	0	5	32	1	16
Martius	2	28	42	30	4	3	56	33	4	9	51	3	0	7	33	1	19
Aprilis	3	28	16	39	5	4	55	42	5	12	19	4	0	10	34	1	21
Maius	4	28	49	58	6	5	54	50	6	14	47	5	0	12	35	1	24
Iunius	5	28	24	7	7	6	53	58	7	17	15	6	0	15	36	1	26
Iulius	6	28	57	26	8	7	53	7	8	19	34	7	0	17	37	1	29
Augustus	7	29	30	44	9	8	52	15	9	22	11	8	0	20	38	1	31
September	8	29	4	54	10	9	51	23	10	24	38	9	0	22	39	1	34
October	9	29	38	12	11	10	50	32	11	27	6	10	0	25	40	1	36
November	10	29	12	22	12	11	49	40	12	29	34	11	0	27	41	1	39
December	11	29	45	40	13	12	48	48	13	32	2	12	0	30	42	1	41
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ANNI BISEXTILIS				15	14	47	5	15	36	58	14	0	34	44	1	46	15
S. G. /				16	15	46	13	16	39	26	15	0	37	45	1	48	20
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Februarius	1	29	8	20	18	17	44	30	18	44	21	17	0	42	47	1	53
Martius	2	29	41	38	19	18	43	38	19	46	49	18	0	44	48	1	56
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Augustus	8	0	29	53	24	23	39	20	24	59	8	23	0	57	53	2	8
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October	10	0	37	21	26	25	37	37				25	1	2	55	2	13
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December	0	0	44	49	28	27	35	53				27	1	7	57	2	18
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					30	29	34	10				29	1	11	59	2	23
					31	30	33	18				30	1	14	60	2	25
IN MENSIBVS				IN DIEBVS				IN HORIS			IN MINVTIS						35

TABVLA

TABVLA PROSTHAPHÆRESIVM SOLARIVM NOSTRIS TEMPORIBVS ACCOMMODATA																																					
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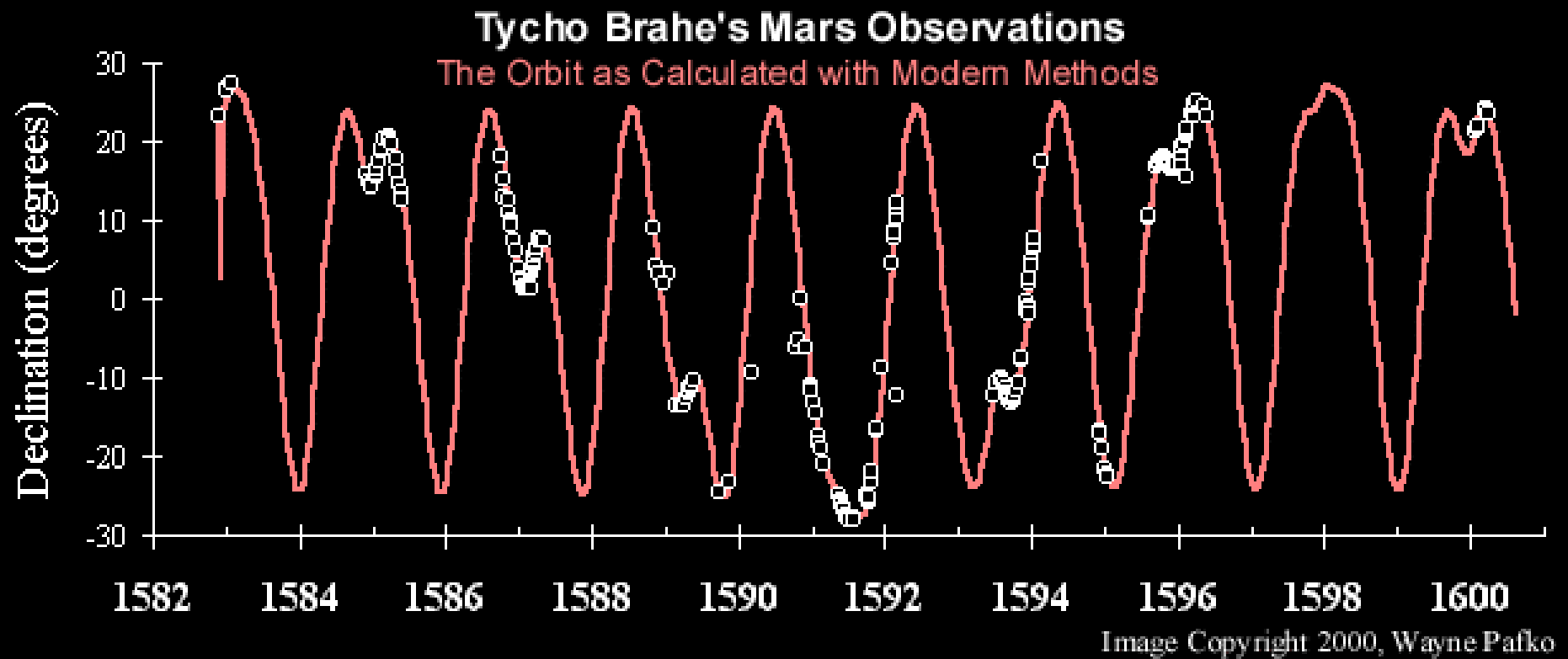
Tycho Brahe's Mars Observations

Source: Tychonis Brahe Dani Opera Omnia

Input by: Wayne Pafko (March 24, 2000)

[MS] = Mars Symbol (you know...the "male" sign)

Year	Month	Day	Day (adj)	Hour	Min	Days since 1 AD	Date	Dec (deg)	Dec (min)	Dec (sec)	Declination
1582	11	12	22			578150.50	1582.89	1.00	23	7	23.12
1582	12	30	40			578198.50	1583.02	1.00	26	56	26.93
1582	12	27	37			578195.50	1583.01	1.00	26	22	26.37
1583	1	18	28			578217.50	1583.07	1.00	27	18	27.30
1584	11	13	23	13	26	578883.06	1584.90	1.00	15	54	15.90
1584	11	27	37	2	15	578896.59	1584.93	1.00	14	42	14.70
1584	12	20	30			578919.50	1585.00	1.00	14	24	14.40
1584	12	21	31			578920.50	1585.00	1.00	14	21	14.36
1584	12	21	31			578920.50	1585.00	1.00	14	21	14.35
1585	1	7	17			578937.50	1585.04	1.00	15	35	15.58
1585	1	9	19			578939.50	1585.05	1.00	15	50	15.83
1585	1	14	24	16	40	578945.19	1585.07	1.00	16	27	16.45
1585	1	22	32	14	55	578953.12	1585.09	1.00	17	31	17.52
1585	1	31	41			578961.50	1585.11	1.00	18	43	18.72
1585	2	3	13	9	43	578964.90	1585.12	1.00	19	1	19.02
1585	2	3	13	9	39	578964.90	1585.12	1.00	19	3	19.05
1585	2	3	13	6	15	578964.76	1585.12	1.00	19	2	19.03
1585	2	4	14	9	14	578965.88	1585.12	1.00	19	9	19.16
1585	2	4	14	8	16	578965.84	1585.12	1.00	19	8	19.13
1585	2	4	14	6	40	578965.78	1585.12	1.00	19	9	19.16
1585	2	17	27	9	45	578978.91	1585.16	1.00	20	21	20.36
1585	2	17	27	9	30	578978.90	1585.16	1.00	20	21	20.36
1585	2	17	27	9	50	578978.91	1585.16	1.00	20	21	20.36
1585	3	12	22	9	20	579001.89	1585.22	1.00	20	32	20.55
1585	3	16	26	7	50	579005.83	1585.23	1.00	20	23	20.38
1585	3	19	29	8	15	579008.84	1585.24	1.00	20	5	20.09
1585	3	26	36	8	20	579015.85	1585.26	1.00	19	44	19.73
1585	4	15	25	9	48	579035.91	1585.31	1.00	17	38	17.64
1585	4	15	25	9	50	579035.91	1585.31	1.00	17	38	17.65
1585	4	26	36	9	50	579046.91	1585.34	1.00	16	8	16.14
1585	5	7	17	11	24.5	579057.98	1585.37	1.00	14	22	14.38
1585	5	7	17	9	20	579057.89	1585.37	1.00	14	22	14.37
1585	5	7	17	11	15	579057.97	1585.37	1.00	14	22	14.38
1585	5	12	22			579062.50	1585.39	1.00	13	30	13.50
1585	5	17	27	11	30	579067.98	1585.40	1.00	12	38	12.64
1585	5	18	28	10	40	579068.94	1585.40	1.00	12	27	12.45



<http://www.pafko.com/tycho>

Johannes Kepler (1571 - 1630)

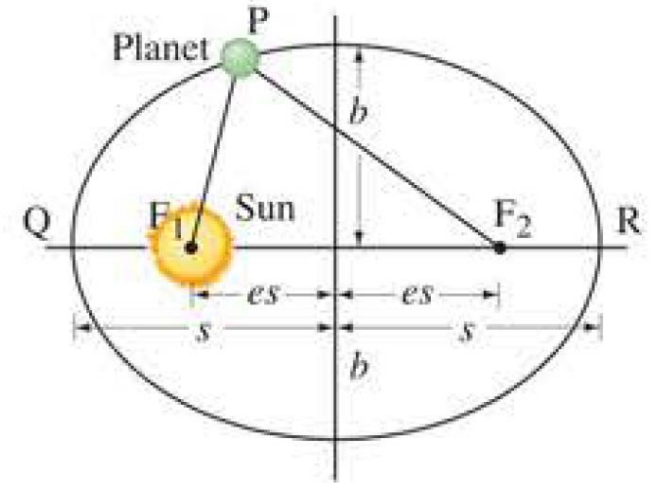
- German astronomer, physicist, and mathematician.
- He was a professor of mathematics in Graz.
- He joined the famous astronomers with the planetary system he explained in his work *Mysterium Cosmographicum* (Mysteries of the Universe, 1596)
- He started working with Tycho Brahe in 1598.
- He made studies on the **movements of the planets** from astronomical observations compiled by Tycho Brahe.
- He changed the face of astronomy with his great contributions.



Kepler's Laws of Planetary Motion

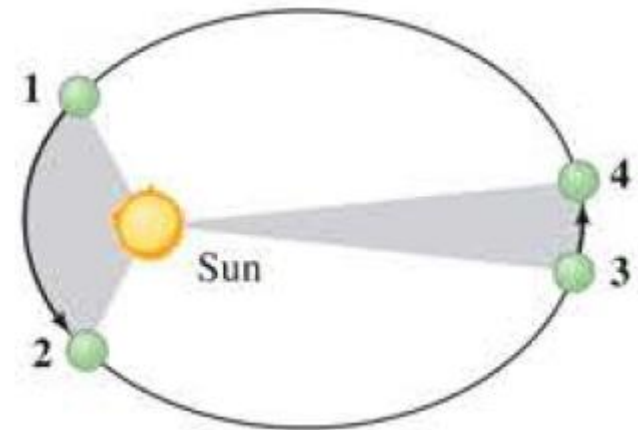
Kepler's first law:

The path of each planet about the Sun is an ellipse with the Sun at one focus.



Kepler's second law:

Each planet moves so that an imaginary line drawn from the Sun to the planet sweeps out equal areas in equal periods of time.



Kepler's third law:

The ratio of the squares of the period a planet to cube of its semimajor axis is a constant (same for all planets)

$$\frac{T^2}{s^3} \equiv \text{constant}$$

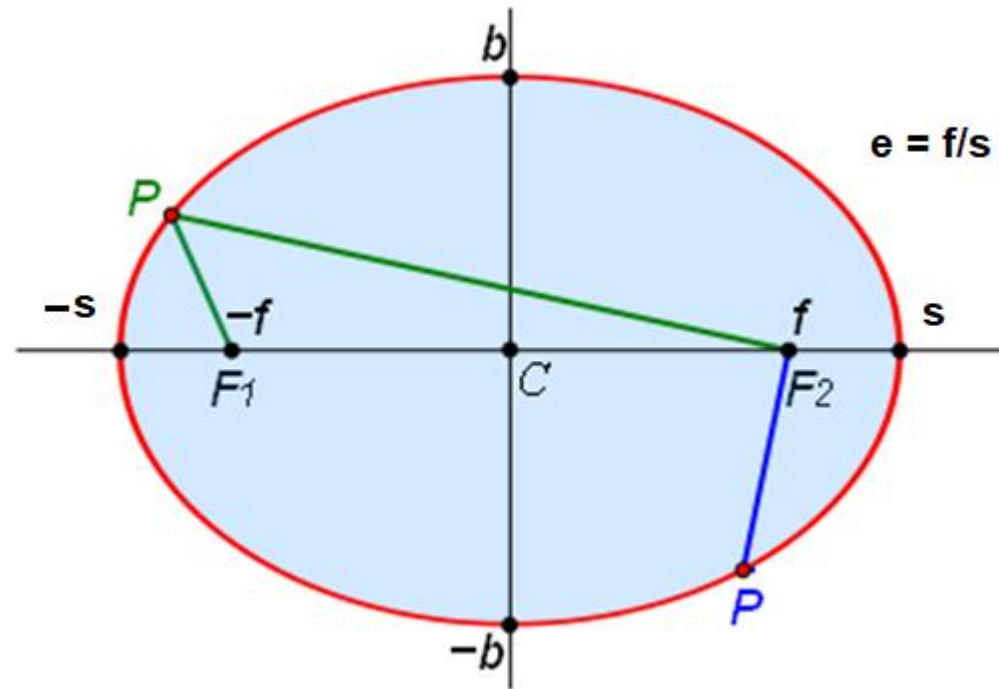
T = period of planet

s = semimajor axes

f = one of the focus of ellipse

$e = f / s$ = eccentricity

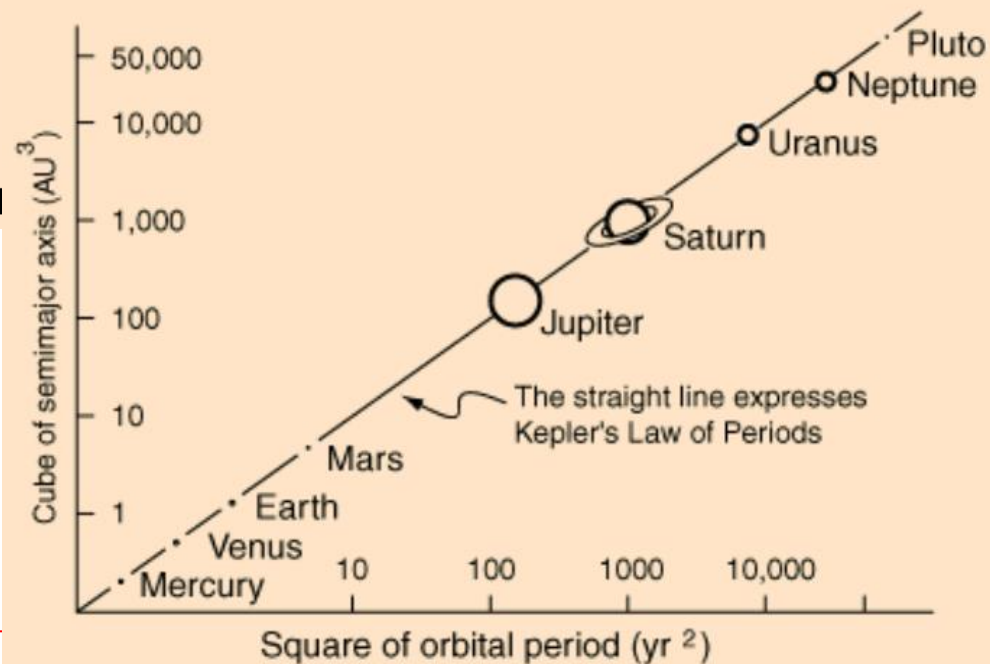
($e = 0$ for circle and $e = 0.0167$ for Earth)



Planet	Semimajor axis (10^{10}m)	Period T (y)	T^2/a^3 ($10^{-34}\text{y}^2/\text{m}^3$)
Mercury	5.79	0.241	2.99
Venus	10.8	0.615	3.00
Earth	15.0	1	2.96
Mars	22.8	1.88	2.98
Jupiter	77.8	11.9	3.01
Saturn	143	29.5	2.98
Uranus	287	84	2.98
Neptune	450	165	2.99
Pluto	590	248	2.99

$$\frac{T^2}{a^3} \equiv \text{constant}$$

Halliday, Resnick, Walker, Fundamentals of Physics 4th ed

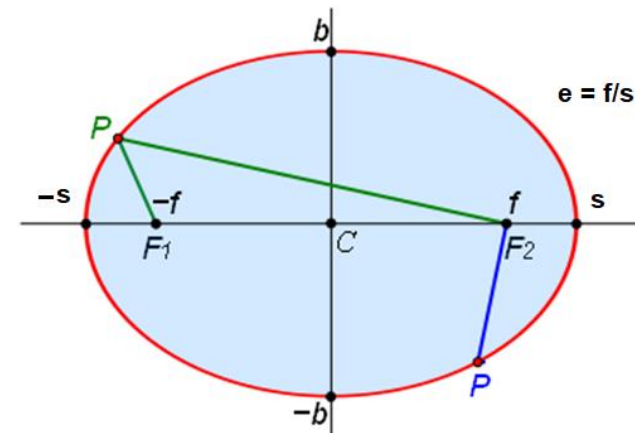
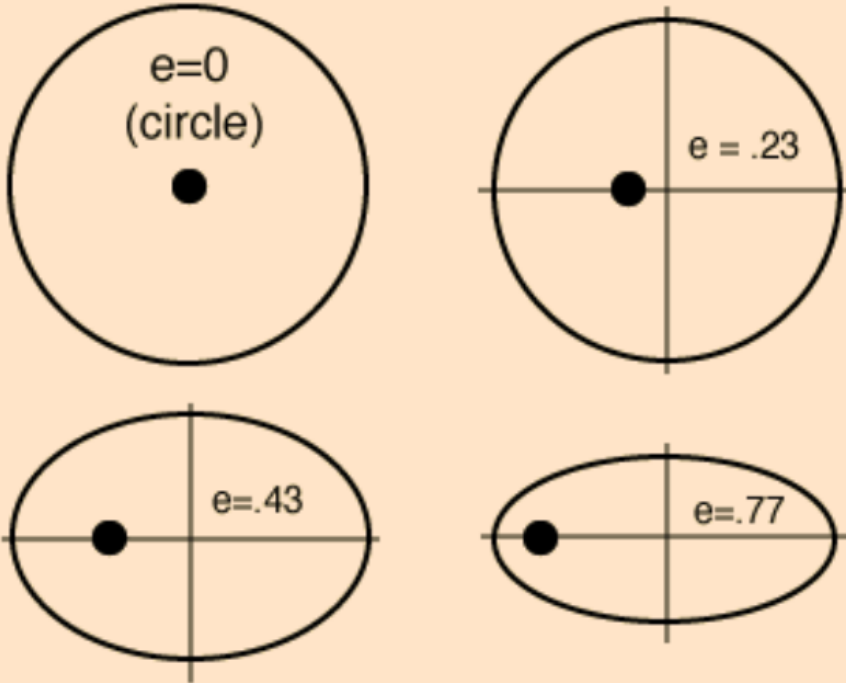


Examples of Ellipse Eccentricity

Planetary orbit
eccentricities

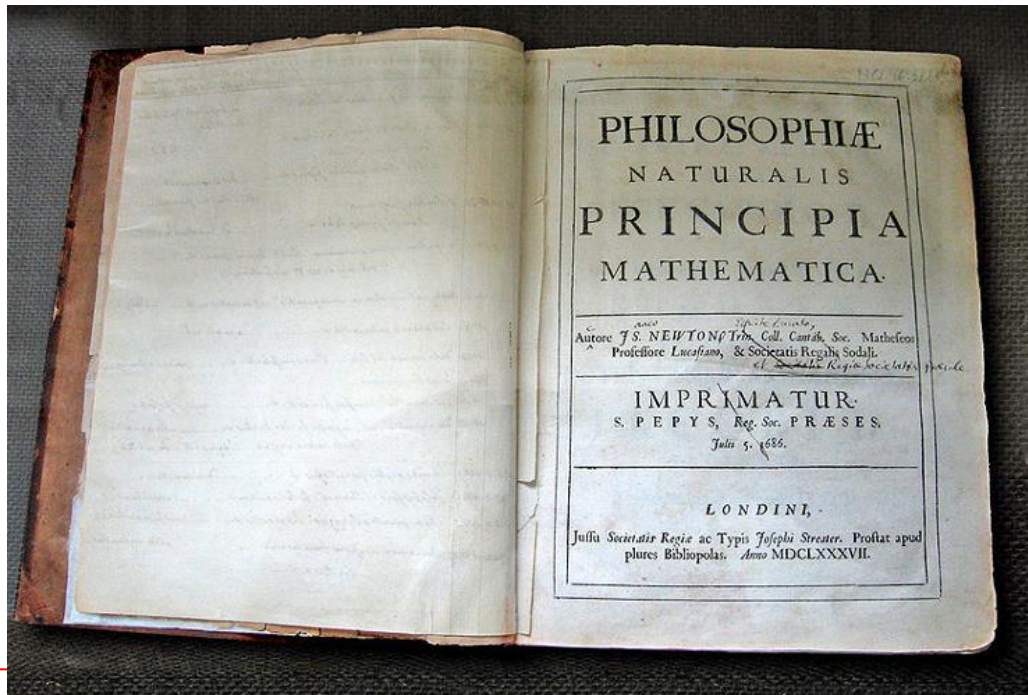
Mercury	0.206
Venus	0.0068
Earth	0.0167
Mars	0.0934
Jupiter	0.0485
Saturn	0.0556
Uranus	0.0472
Neptune	0.0086
Pluto	0.25

$$e = \frac{f}{s}$$



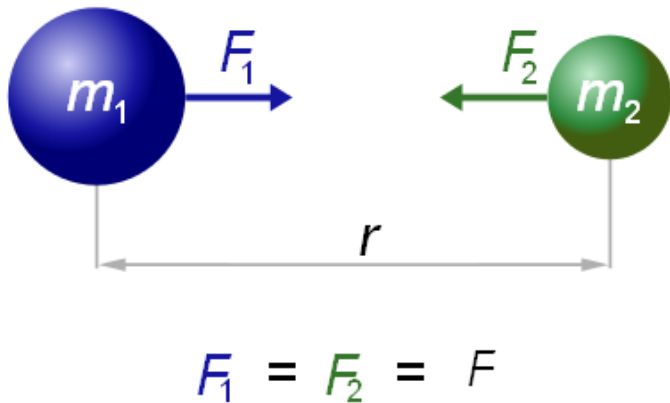
Newton's Law of Universal Gravitation

- In 1687, Newton published his famous work Mathematical Principles of Natural Philosophy.
- In this book, he mentioned:
Laws of Motion and **Law of Universal Gravitation**
- You can find the book at:
<https://archive.org/details/philosophiaenatu00newt/page/n5/mode/2up>



Law of universal gravitation:

Every particle in the universe attracts every other particle with a force that is proportional to the product of their masses and inversely proportional to the square of the distance between them. This force acts along the line joining the two particles.



$$F = G \frac{m_1 m_2}{r^2}$$

m_1, m_2 are masses of particles

r is distance between them

G is Universal Gravitational Constant

$$G = 6.673 \times 10^{-11} \text{ N.m}^2/\text{kg}^2$$

Measurement of Constant G

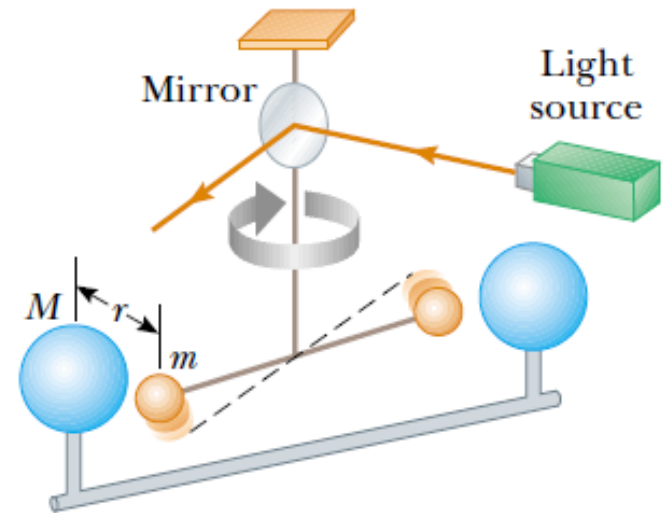
The Universal Gravitational Constant was first measured by H. Cavendish in 1798.

In the experiment, two small and two large masses are used.

G value is evaluated by measuring the deviation angle between incoming and reflected light rays as shown in Figure.

See:

http://en.wikipedia.org/wiki/Cavendish_experiment



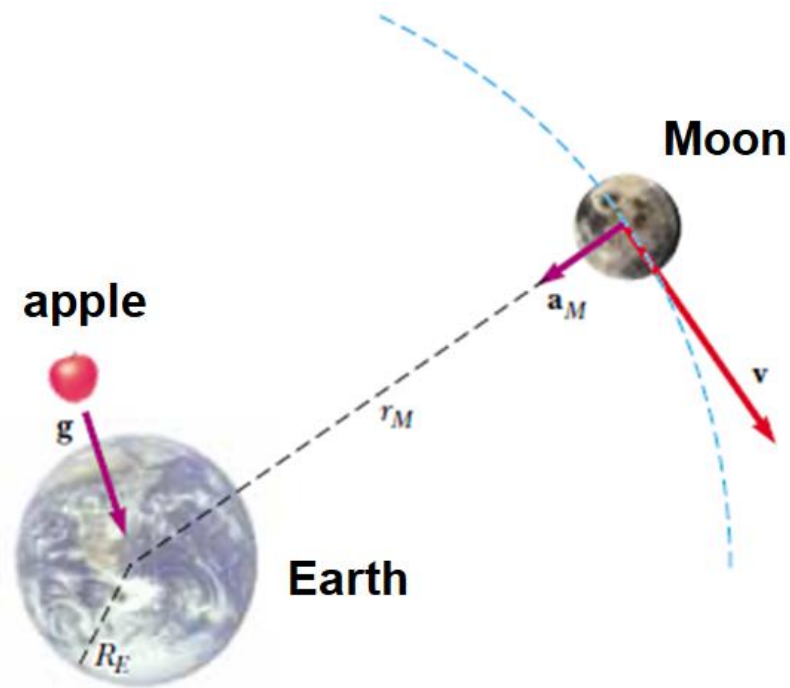
Example:

What is the gravitational force between two people each weighing 75 kg if the distance between them is half a meter?

$$F = (6.673 \times 10^{-11} \text{ N.m}^2/\text{kg}^2) \frac{(75 \text{ kg})(75 \text{ kg})}{(0.5 \text{ m})^2} = 1.5 \times 10^{-6} \text{ N}$$

How about moon?

- Newton compared the acceleration of the Moon and that of a freely falling body on the surface of the earth.
- If the source of both accelerations is the same (*we do'nt know mass of earth*)



$$\frac{\text{Acceleration of moon}}{\text{Acceleration of apple}} = \frac{a_M}{g} = \frac{GM_E/r_M^2}{GM_E/R_E^2} = \left(\frac{R_E}{r_M}\right)^2 = \left(\frac{6.37 \times 10^6 m}{3.84 \times 10^8 m}\right)^2 = 2.75 \times 10^{-4}$$

- Hence, acceleration of the moon

$$a_M = (2.75 \times 10^{-4})(9.8 \text{ m/s}^2) = 2.70 \times 10^{-3} \text{ m/s}^2$$

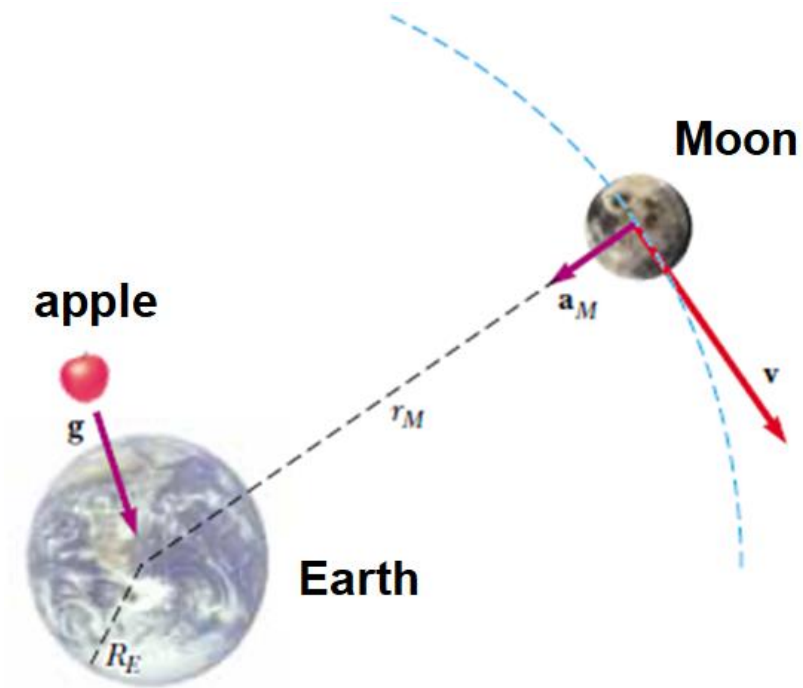
- On the other hand

$$a_M = \frac{v^2}{r_M} = 2.72 \times 10^{-3} \text{ m/s}^2$$

- Percentage Error:

$$P.E. = \frac{2.72 \times 10^{-3} - 2.70 \times 10^{-3}}{2.70 \times 10^{-3}} \times 100 = \% 0.7$$

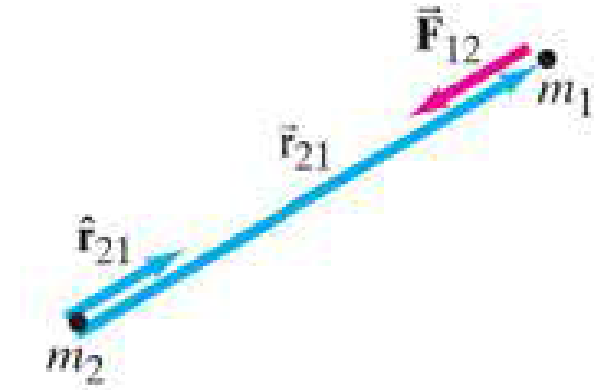
- This result strongly confirms the inverse-square law.



Vector Form of Newton's Law of Gravitation

We can write Newton's law of universal gravitation in vector form as

$$\mathbf{F}_{12} = -G \frac{m_1 m_2}{r_{21}^2} \hat{\mathbf{r}}_{21}$$



When many particles interact, the total gravitational force on a given particle is the vector sum of the forces exerted by others.

$$\mathbf{F}_1 = \mathbf{F}_{12} + \mathbf{F}_{13} + \cdots + \mathbf{F}_{1n}$$

Gravity Near the Earth's Surface

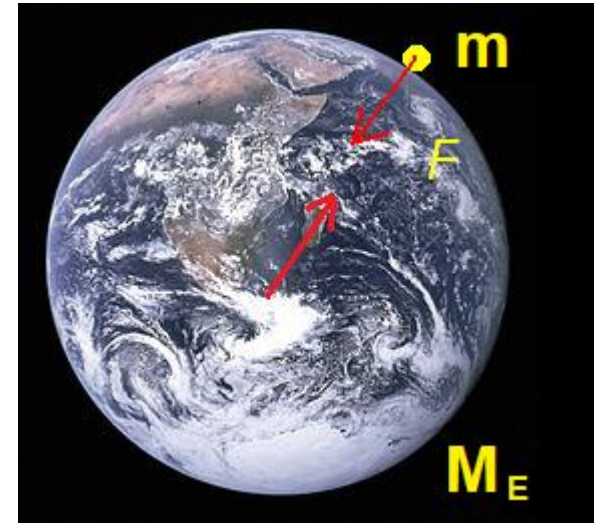
Formula can be applied to objects having spherical symmetry.

So, gravitational force and acceleration of an object on Earth surface is:

$$F = G \frac{M_E m}{R_E^2} = mg \quad \Rightarrow \quad g = G \frac{M_E}{R_E^2}$$

Solving for **Mass of Earth**:

$$M_E = \frac{g R_E^2}{G} = \frac{(9.8 \text{ kg.m/s}^2)(6.38 \times 10^6 \text{ m})^2}{6.673 \times 10^{-11} \text{ N.m}^2/\text{kg}^2} = 6 \times 10^{24} \text{ kg}$$



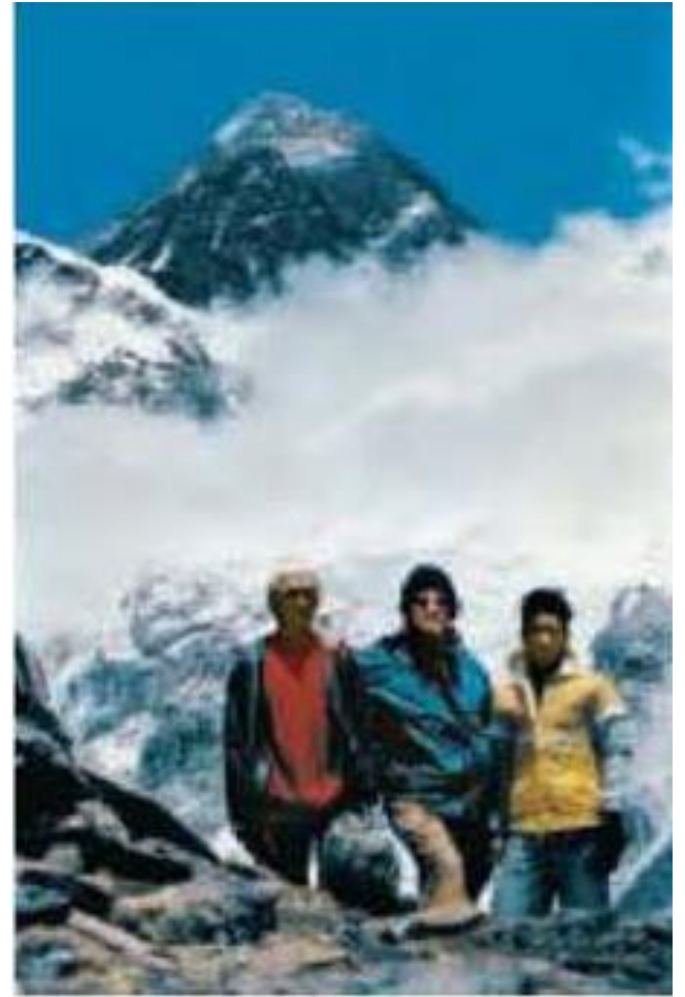
Example

Estimate the effective value of g on the top of Mt. Everest, 8850 m above sea level .

Solution

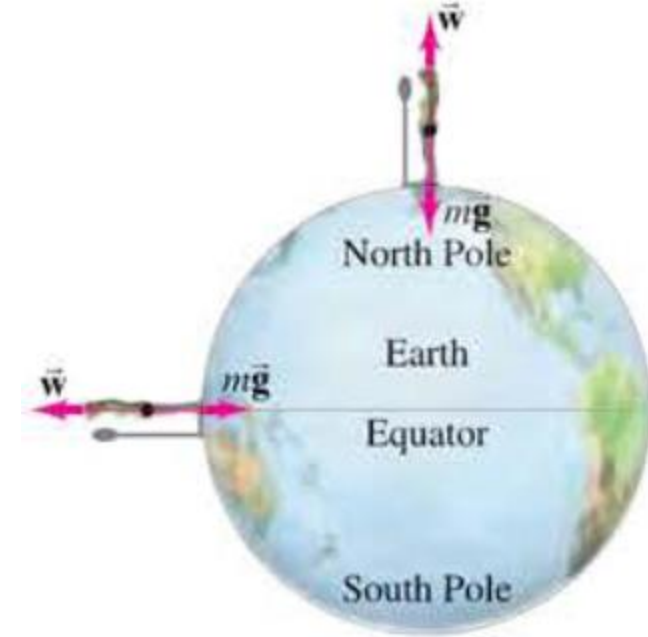
$$r = 6380 \text{ km} + 8.85 \text{ km} = 6388.9 \text{ km}$$

$$g = G \frac{M_E}{r^2} = 9.77 \text{ m/s}^2$$



Example: Effect of Earth's rotation on g.

Assuming the Earth is a perfect sphere, determine how the Earth's rotation affects the value of g at the equator compared to its value at the poles.



Solution

North pole: Normal force – gravitational force = 0 ($w - mg = 0$)

Equator:

Speed of an object at rest on equator: $v = \frac{2\pi R_E}{1 \text{ day}} = 4.64 \times 10^2 \text{ m/s}$

Consider centripetal force: $w - mg' = \frac{mv^2}{R_E}$

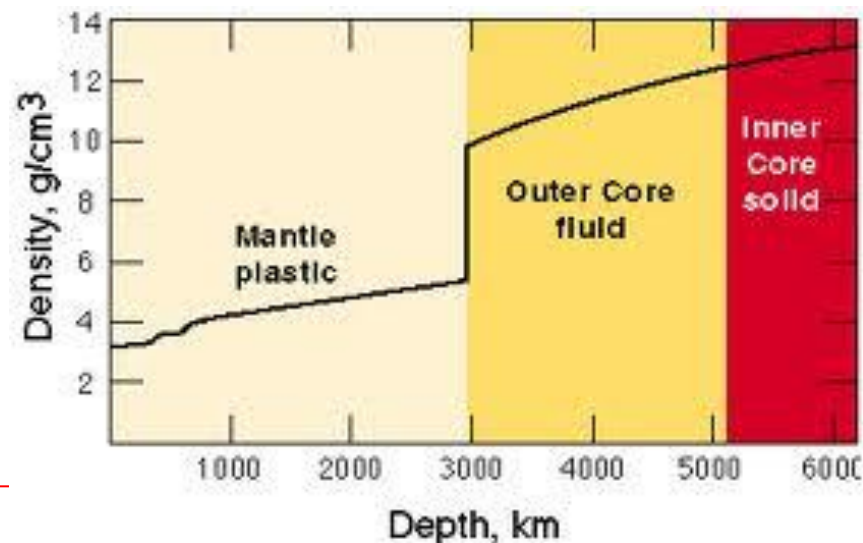
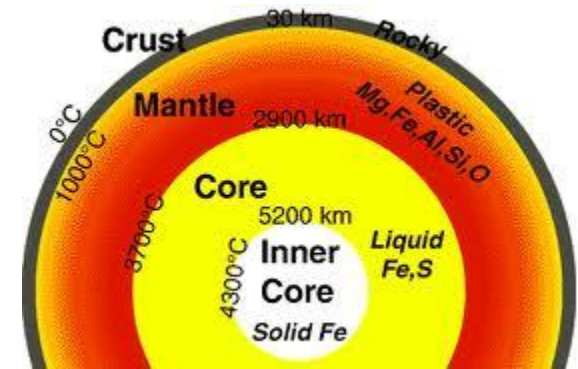
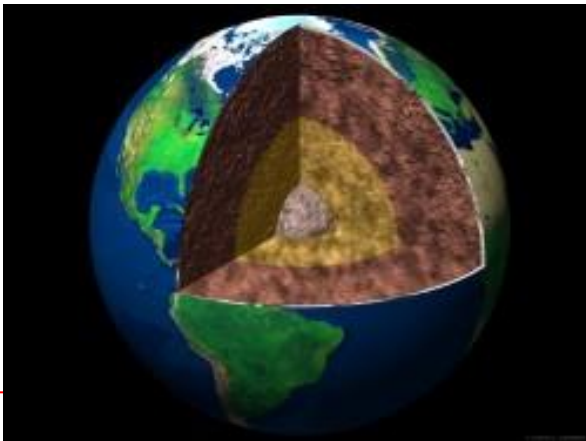
$$\text{Or } g' = \frac{w}{m} - \frac{mv^2}{mR_E} = \frac{mg}{m} - \frac{mv^2}{mR_E} = g - \frac{v^2}{R_E} \Rightarrow \Delta g = \frac{v^2}{R_E} = 0.0337 \text{ m/s}^2$$

Example: Density of Earth

$$\text{Average Density} = \frac{\text{mass}}{\text{volume}} = \frac{6 \times 10^{24} \text{ kg}}{\frac{4}{3} \pi (6.4 \times 10^6 \text{ m})^3} = 5.5 \text{ g/cm}^3$$

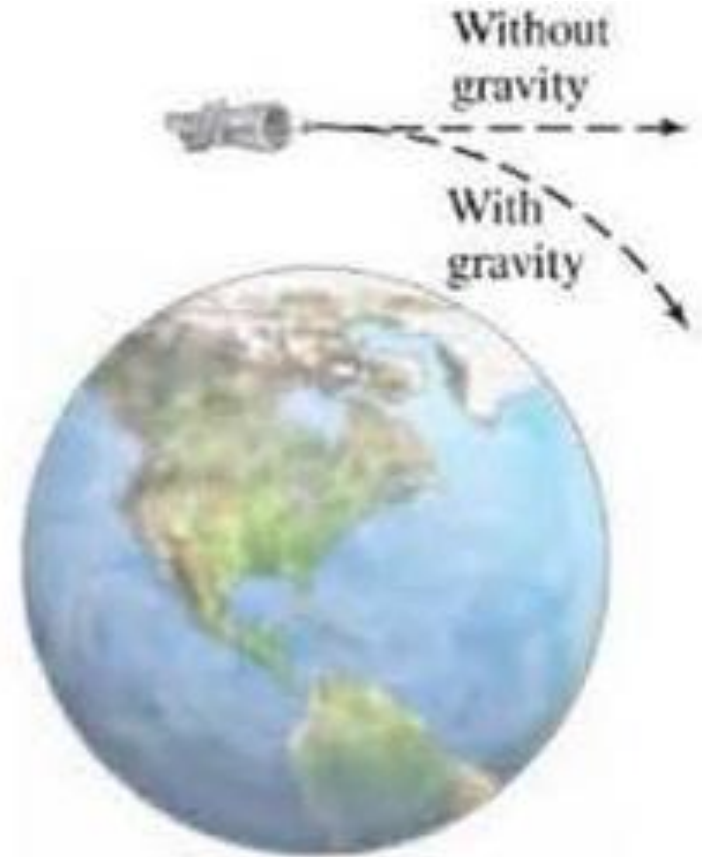
Density of objects (rock, soil, etc) on Earth surface is about **3 g/cm³**.

Therefore the density of the earth's interior is much higher!



Satellite Motion

What keeps a satellite up? => its speed and gravity



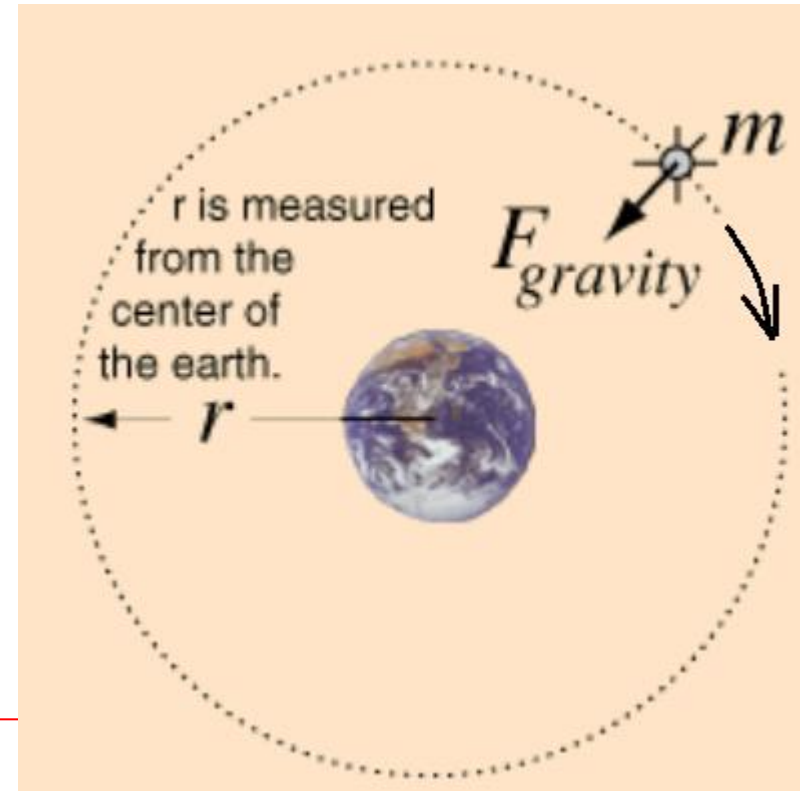
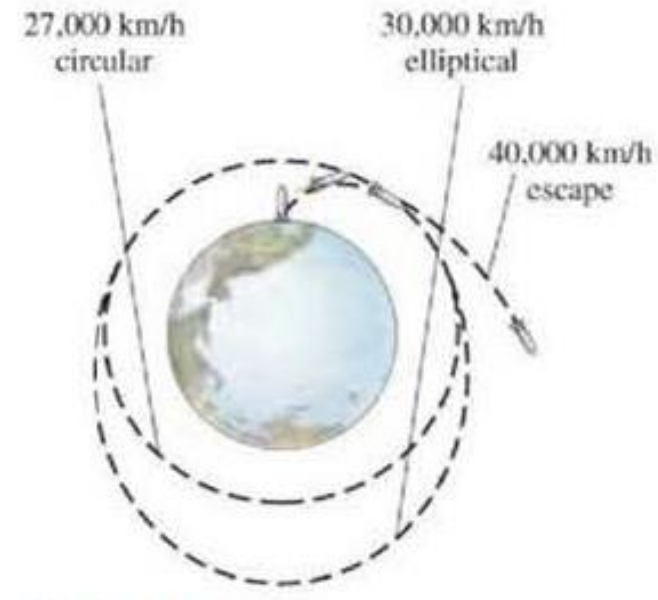
Satellite Motion

Force on satellite of mass m
in circular orbit:

$$F = G \frac{mM_E}{r^2} = \frac{mv^2}{r}$$

Solving for velocity:

$$v = \sqrt{\frac{GM_E}{r}}$$



Example:

Find the height above the earth surface and velocity of a satellite whose period is 1 day.

Solution

Speed of the satellite: $v = \frac{2\pi r}{T}$ and $v = \sqrt{\frac{GM_E}{r}}$

Equating them:

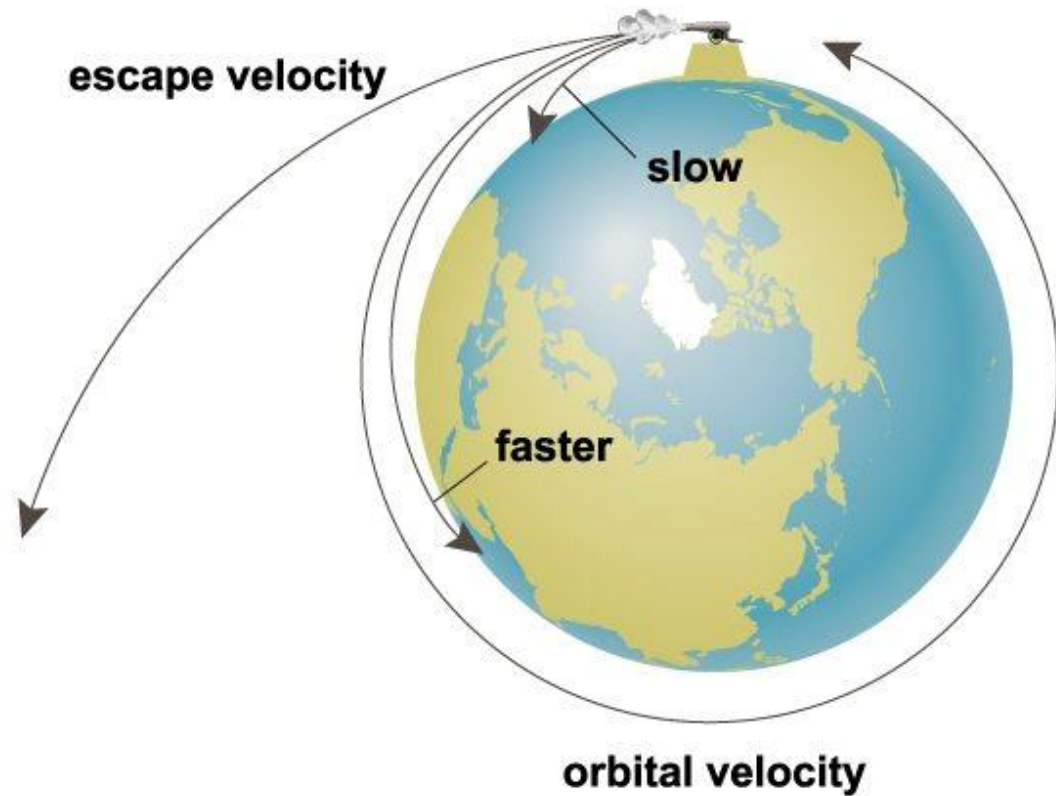
$$\frac{2\pi r}{T} = \sqrt{\frac{GM_E}{r}}$$

Solving for radius: $r = 42000$ km from center earth.

Solving for the speed: $v = 3070$ m / s

Example:

Calculate the critical speed on the Earth surface for a circular orbit.



$$v = \sqrt{\frac{GM_E}{R_E}} \approx 7900 \text{ m/s}$$

Weightlessness

In elevator, apply Newton's 2nd law:

in general

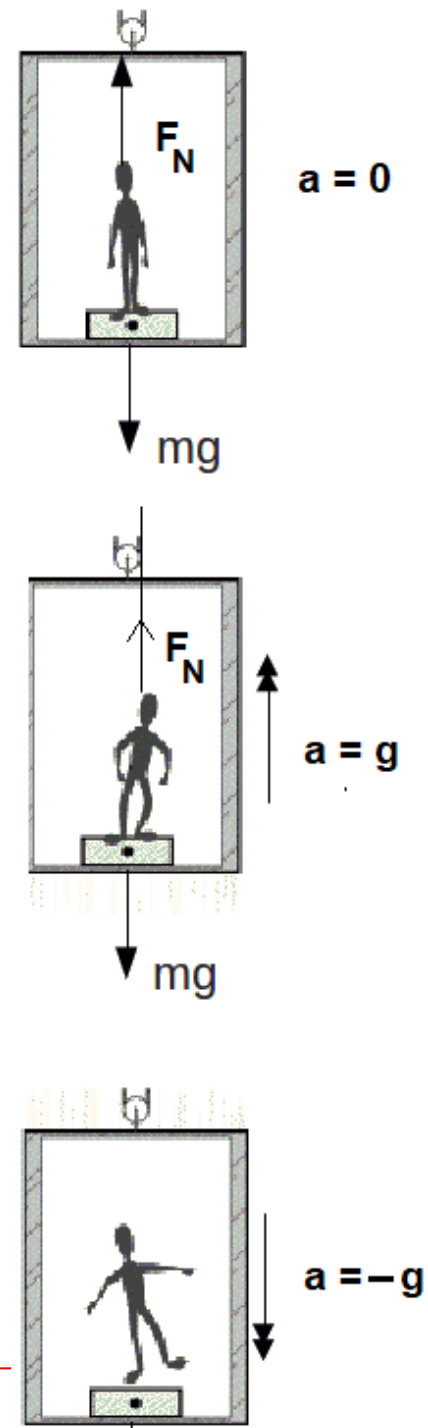
$$F_N - mg = ma$$

$$a = 0 \Rightarrow F_N - mg = 0 \Rightarrow F_N = mg$$

$$a = g \Rightarrow F_N - mg = mg \Rightarrow F_N = 2mg$$

$$a = -g \Rightarrow F_N - mg = -mg \Rightarrow F_N = 0$$

In last condition, you'll feel **weightless**!



The "weightlessness" experienced by people in a satellite orbit close to the Earth is the same apparent weightlessness experienced in a freely falling elevator.

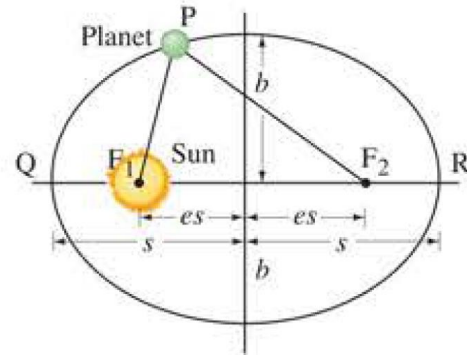


Kepler's Laws and Newton's Synthesis

Kepler's 1st and 2nd laws require more analysis.

Kepler's 3rd law:

$$\frac{T^2}{s^3} \equiv \text{const}$$



Assuming, nearly circular orbit and put average distance $s = r$,

$$F = G \frac{M_E M_S}{r^2} = \frac{M_E v^2}{r} \quad \text{and} \quad v = \frac{2\pi r}{T} \Rightarrow \frac{T^2}{r^3} = \frac{4\pi^2}{GM_S}$$

*Using this relation,
we can evaluate the mass of the sun!*

How?

Example:

Period of Earth: $T = 1$ year,
Earth-Sun distance: $r = 1.5 \times 10^{11}$ m,
Evaluate mass of Sun.

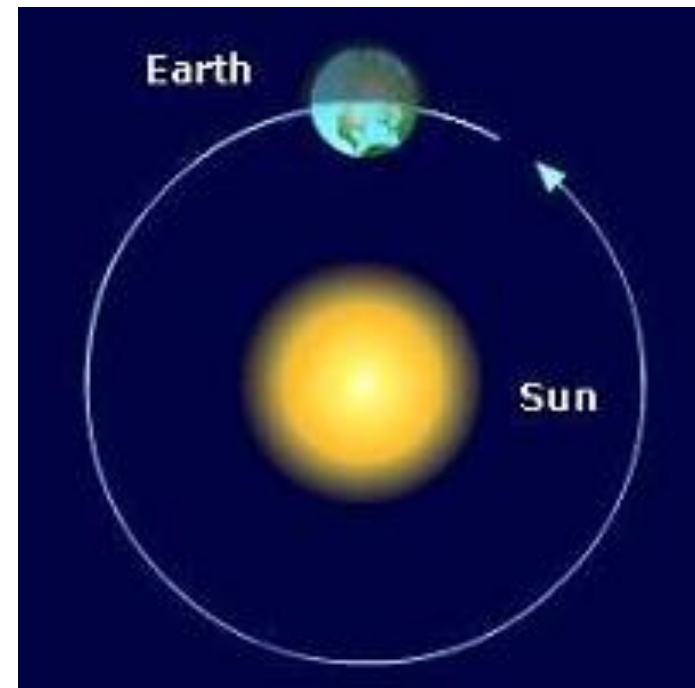
Solution:

From last relation:

$$M_S = \frac{4\pi^2 r^3}{GT^2} = 2 \times 10^{30} \text{ kg}$$

Q: Can you find mass of Earth using Moon's data?

A: Yes. It gives $M_E = 6 \times 10^{24}$ kg.



For any two planets:

$$\frac{T_1^2}{r_1^3} = \frac{T_2^2}{r_2^3} = \text{const} = \frac{4\pi^2}{GM_S}$$

Example:

Determine the mean distance of Mars from the Sun using the Earth as a reference. ($T_{\text{Mars}} = 687$ days)

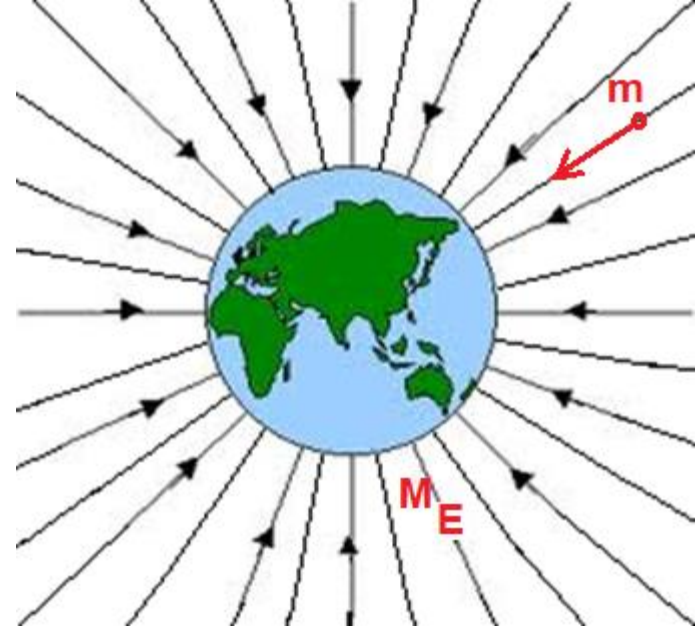
$$\frac{(365)^2}{1^3} = \frac{(687)^2}{r_2^3} \Rightarrow r_2 = 1.52$$

*So, Mars is 1.52 times the Earth's distance from the Sun,
or 2.28×10^{11} m.*

Gravitational Field

Most of the forces we meet in everyday life are contact forces: **you push or pull something.**

The gravitational force acts over a distance.
Earth exerts a force on a falling apple.



We can define the gravitational field as
the gravitational force per unit mass at any point in space.

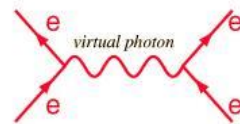
Near Earth surface at a distance from center, we can evaluate field as:

$$\vec{g} = \frac{\vec{F}}{m} = -\frac{GmM_E}{mr^2} \hat{r} = -\frac{GM_E}{r^2} \hat{r}$$

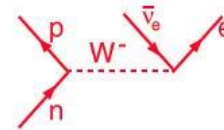
$\hat{r}$ = the unit vector to the direction of center (shown as red in figure).

Four Fundamental Forces (Interactions)

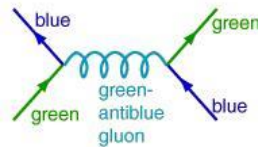
Force	Source	Range (m)	Intensity	Theory	Particle
Gravity	Mass	∞	1	General Relativity	Graviton ?
Weak	Weak charge	10^{-17}	10^{25}	Electroweak	W ve Z
Electro-magnetic	Electric charge	∞	10^{36}	QED	Photon
Strong	Color charge	10^{-15}	10^{38}	QCD	Gluon



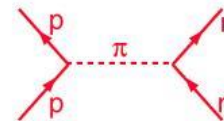
Electromagnetic



Weak



between quarks



between nucleons

Strong Interaction

Dark Matter

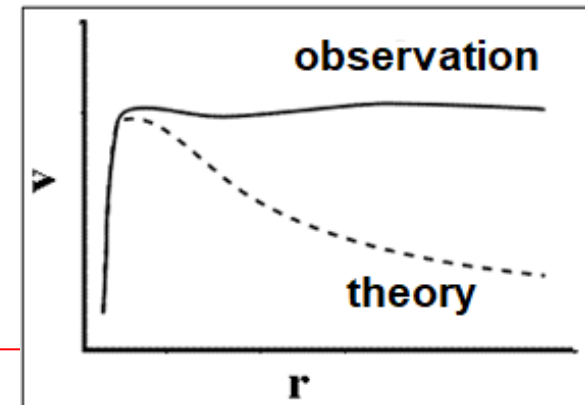
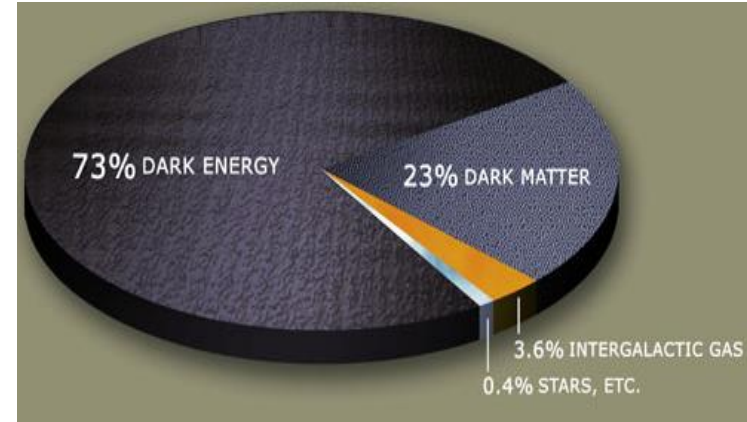
- It is not radiate or reflect light (invisible).
- It interacts with visible matter.
- It surrounds the whole Galaxy!

Evidence of dark matter:

$$\frac{mv^2}{r} = G \frac{mM}{r^2} \rightarrow v = (GM / r)^{1/2}$$

Center of galaxy: $M \propto r^3 \rightarrow v \propto r$

Outer part: $M \cong \text{constant} \rightarrow v \propto r^{-1/2}$





Turkish Appendix

Kütleçekimin Doğası

- Kütleçekim yasası çok sade!
- Kütleçekimin arkasındaki işleyiş (mekanizma), Fizik ve Hikmet nedir?
- Neden kütleler birbirini çekiyor?
- Newton bu mekanizma için bir şey söylememiş.
- Fiziğin birçok yasası matematiksel kurallardan oluşur ve mekanizma konusunda bir fikir vermez.
- Neden doğa olaylarını açıklamak için matematik kullanıyoruz ama arkadaki mekanizma ile uğraşmıyoruz? Kimse Bilmiyor.
- Ancak buna devam ediyoruz çünkü işe yarıyor.

K tle ekimin Doęası

K tle ekim mekanizması i in bir  ok fikir  ne s r lm ş.

K   k par acıklar hipotezi (1750)

Uzayın her y ne  ok hızlı hareket eden k   k par acıklarla dolu olduęunu varsayalım. Bu par acıklar D nya ile  arpıřtıęında ona bir itme kazandırır. G neř doęrultusunda gelen par acıklar kısmen daha azdır.   nk  bu par acıkların bir kısmını G neř soęurur. Bu sayede D nya, dięer y nlerden gelen itmelerle G neře doęru s r klenir.

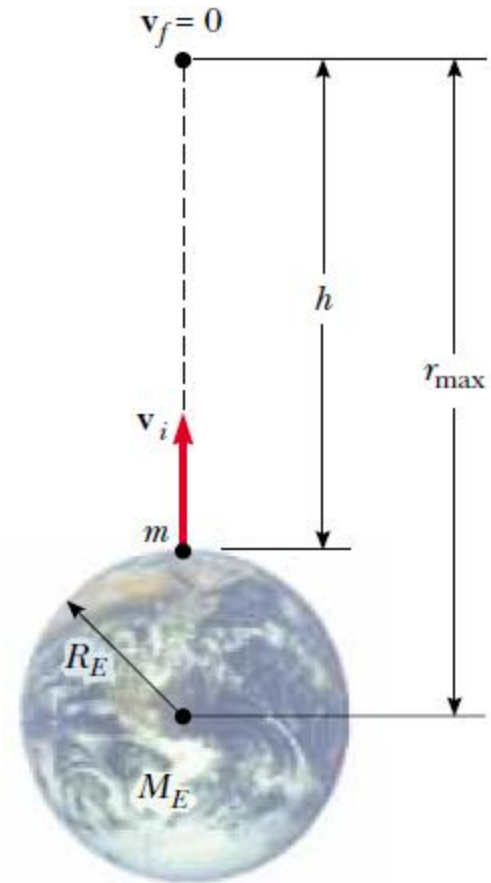
Bu mekanizma   ker!

  nk  (yaęmurda kořan adam misali) D nya G neř etrafında d nerken  n y zey daha  ok itmeye maruz kalır. Bu, d nyanın yavařlamasına sebep olur.

Kurtulma Hızı

Dünya yüzeyinden v hızıyla bir cisim fırlatılsın.
Bu cismin dünyanın çekim etkisinden kurtulabilmesi için gereken en küçük hız:

$$v_{kurtulma} = \sqrt{\frac{2GM}{R}}$$



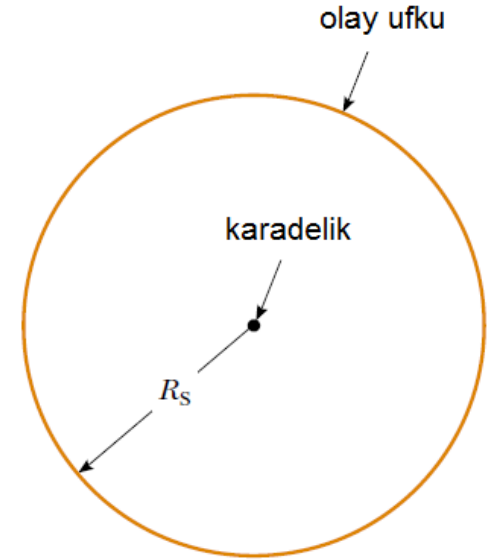
Astronomik nesne	$v_{kurtulma}$
-----	-----
Ay	2.3 km/s
Dünya	11.2 km/s
Jupiter	60.0 km/s
Güneş	618.0 km/s

Karadelikler

Yıldızlar iç yakıtlarını tükettince, kütleçekimi yıldızın içine çökmesine sebep olur.

M = 1.4 Güneş kütlesi üzerinde ise, çökme devam eder => Nötron yıldızı oluşur.

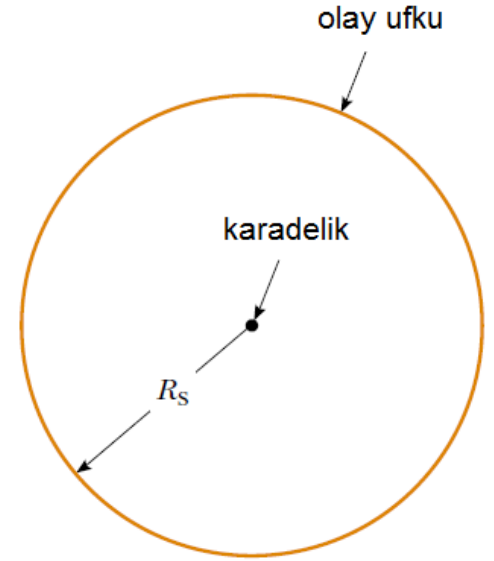
M = 3 Güneş kütlesi üzerinde ise, çökme devam eder => Karadelik oluşabilir.



Karadelikler

Karadelikler için kurtulma hızı çok büyüktür
(Neden?)

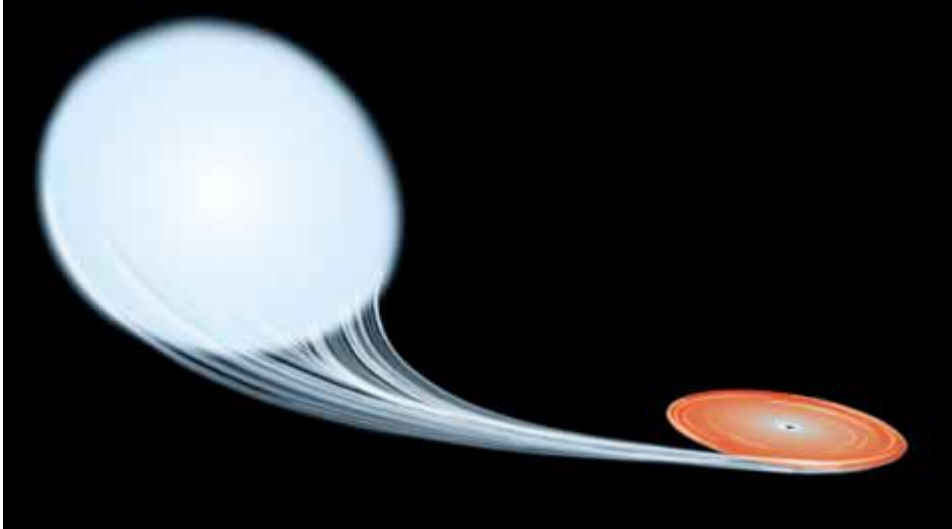
$V_{\text{kurtulma}} = c$ olduğu duruma karşılık gelen
yarıçapa **Schwarzschild yarıçapı** (R_s),
karadeliği çevreleyen R_s yarıçaplı
hayali küreye **olay ufku** denir.



Olay ufku civarında (dışında) hayatta kalmak olasıdır.

Karadelikler

Çift yıldızlardan birinin karadelik olması durumuna ait bir benzetim (simülasyon).



Karadelik etrafındaki madde,
karakteristik x-ışını yaymaya başlar.
Neden?

Genel Görelilik ve Kütleçekim

Newton: “kütleçekim sonsuz hızda iletilir”.

Einstein: “hiçbirşey ışıktan hızlı gidemez”.

Einstein, Görelilik İlkesini dikkate alarak, Kütleçekim yasasında bir düzeltme yapıyor.

Yeni yasa:

Genel Görelilik Yasası

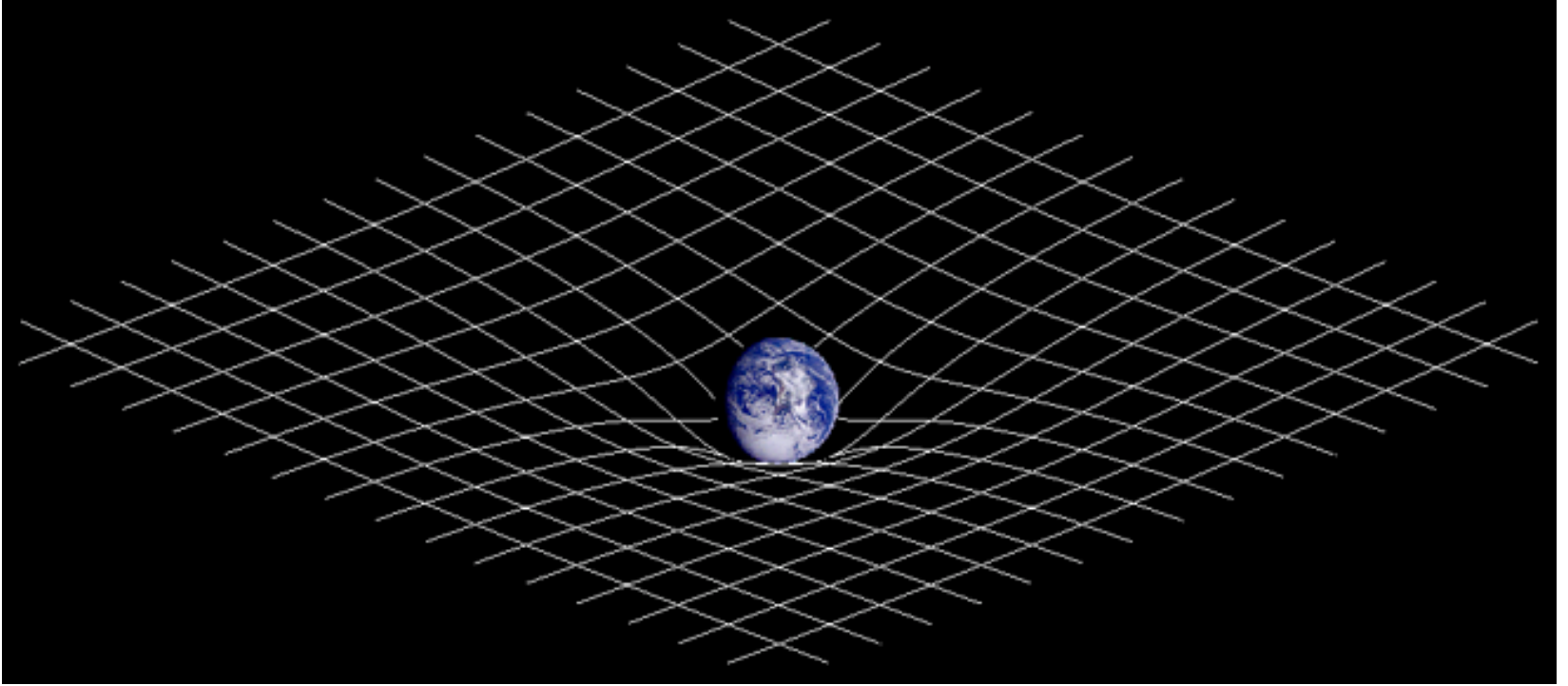
Genel Görelilik ve Kütleçekim

Dahası ($E = mc^2$ den), kütle ve enerji eşdeğerdir ve her ikisi de kütleçekiminden etkilenir.

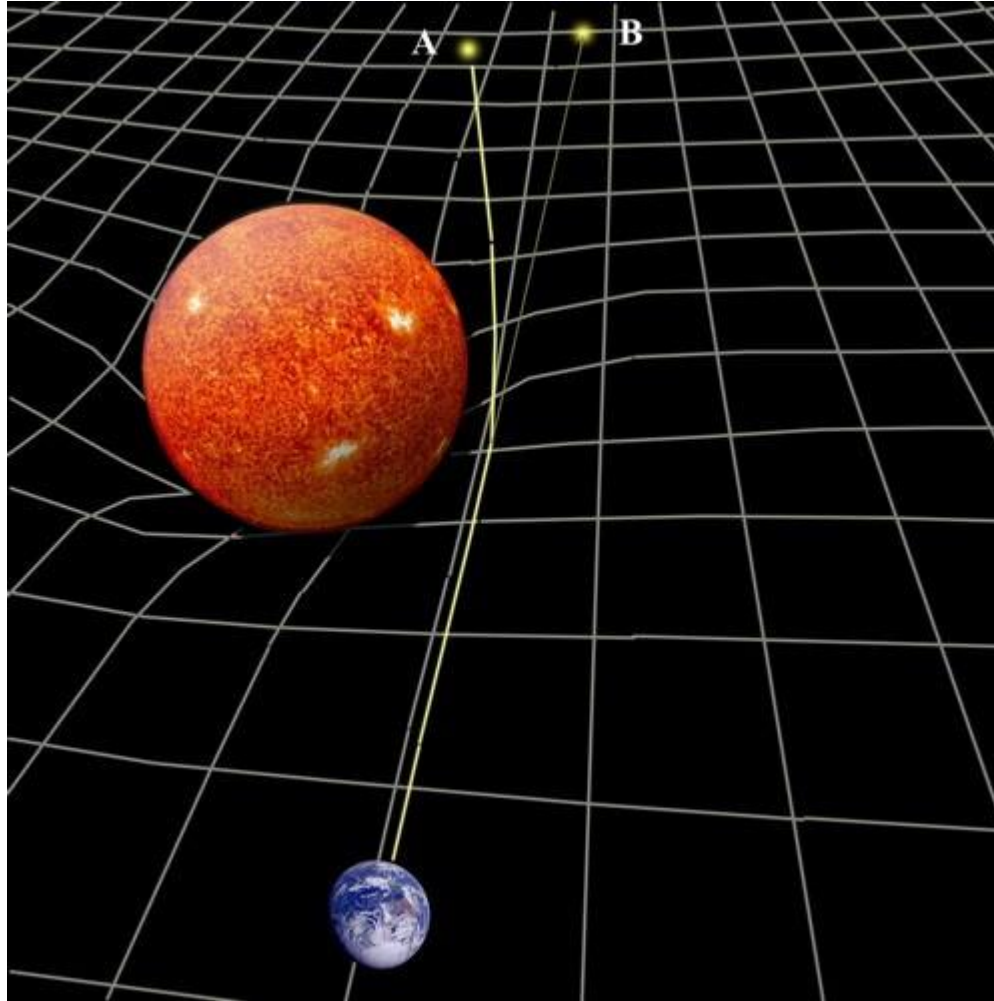
Buna göre, ışık (külliye enerji) uzayda düz dizgi boyunca değil eğri-büğrü bir yol izler.

1. Kanıt: Güneş tutlulması esnasında bir yıldızın konumu
2. Kanıt: Mekür gezegeninin güneş çevresindeki hareketi
3. Kanıt: Kütleçekimsel Mercek etkisi

Genel Görelilik ve Kütleçekim

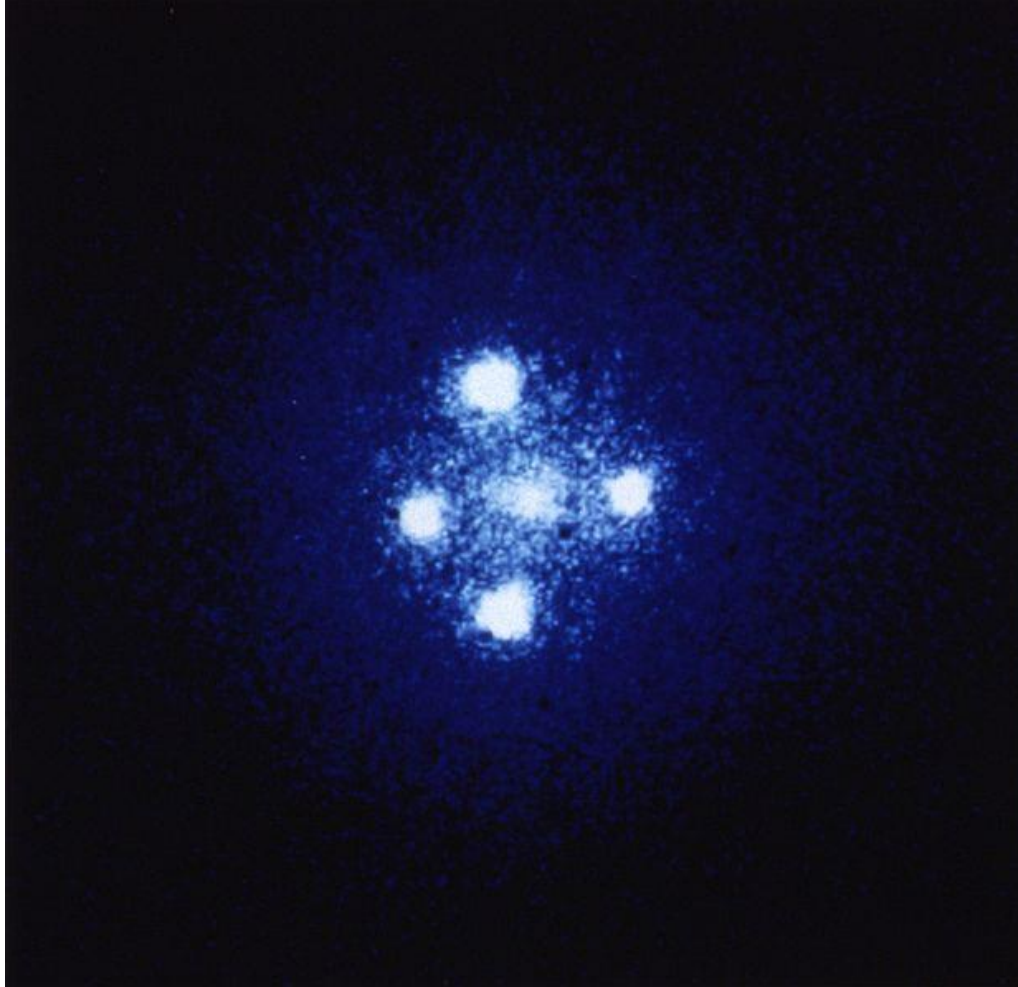


Genel Görelilik ve Kütleçekim



Genel Görelilik ve Kütleçekim

Aynı astronomik nesnenin
kütleçekimsel mercek etkisi ile dört görüntüsü

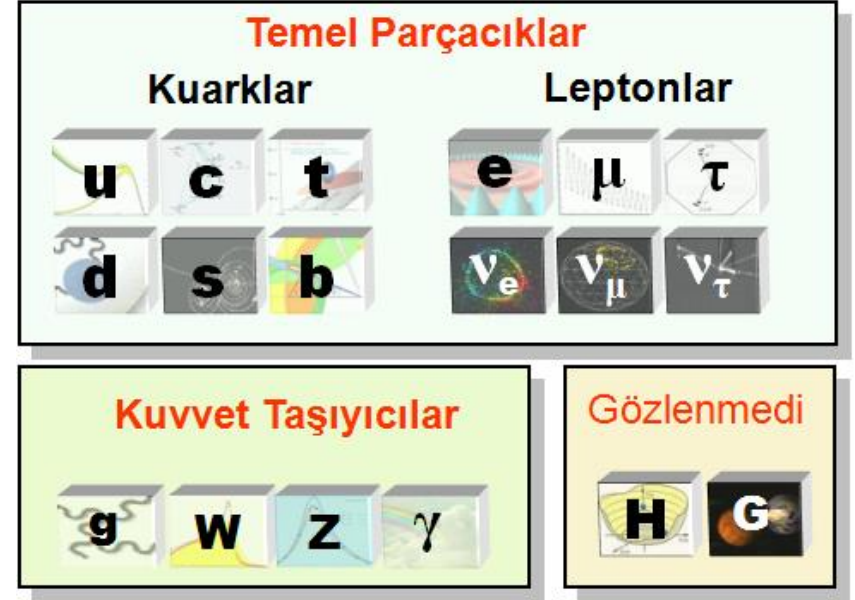


Kuantum Fiziği ve Kütleçekim

Bütün madde temel parçacıklardan oluşmuştur.

Ölçek küçüldükçe,
Kuantum etkileri ortaya çıkar.

Çok zayıf olan Kütleçekim,
henüz Kuantum doğası
kazanmamıştır.



Fizik yasalarının matematiksel olarak tutarlı olmalı.
Kütleçekimin de kuantum yapısı olmalıdır.

Beklenti:

Newton yasası -> Einstein yasası

Einstein yasası -> Kuantum kütleçekimi (belirsizlik ilkesi)

Sicim Kuramı (Herşeyin Kuramı)

String Theory (Theory of Everything)

Amaç:

- Bütün parçacıkları ve dört temel kuvveti bir çatı altında toplamak.

Fikir:

- Temel parçacıklar nokta şeklinde değil titreşen sicim ilmikleridir
- Kuram ek boyutların varlığını öngörüyor:
 $4+6 = 10$ boyut veya $4+22 = 26$ boyut !

